

Illuminating the architecture of the greater Mount St. Helens magmatic system from slab to surface



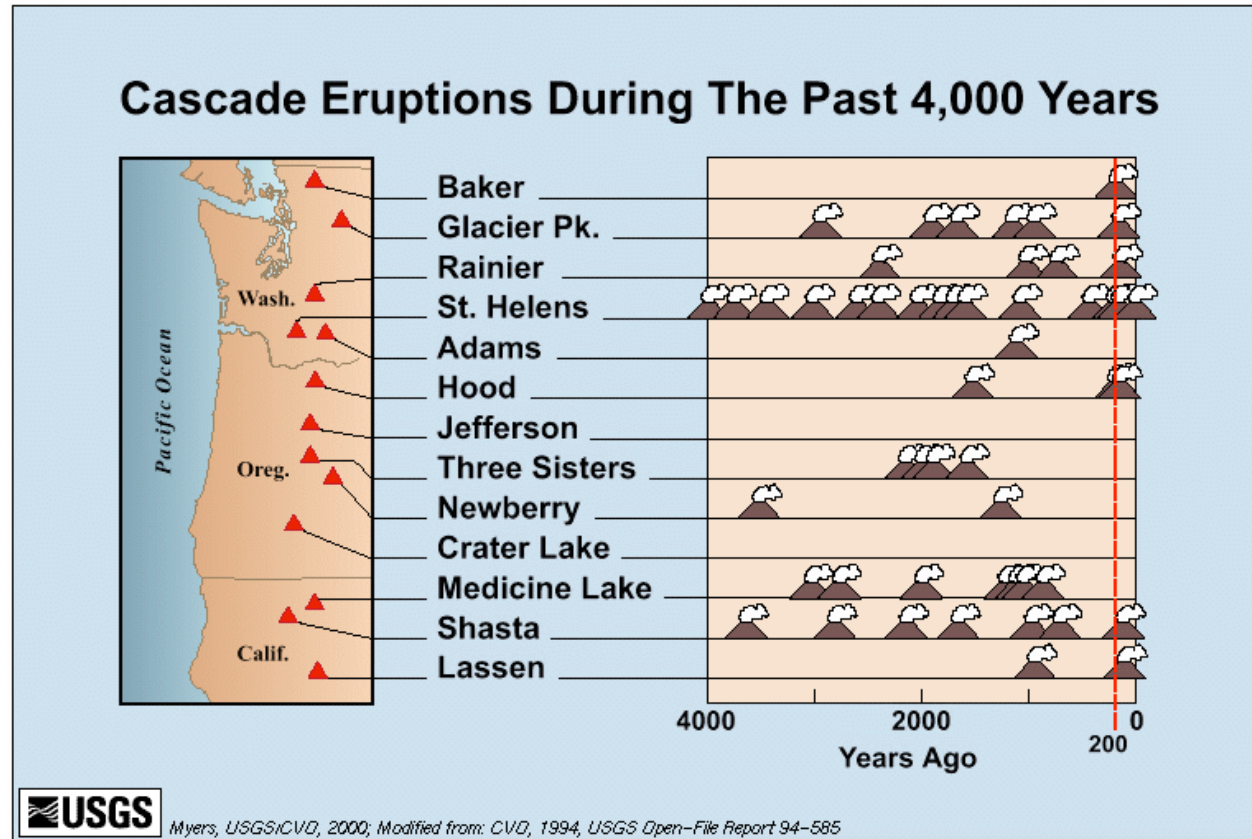


Mount St. Helens Team

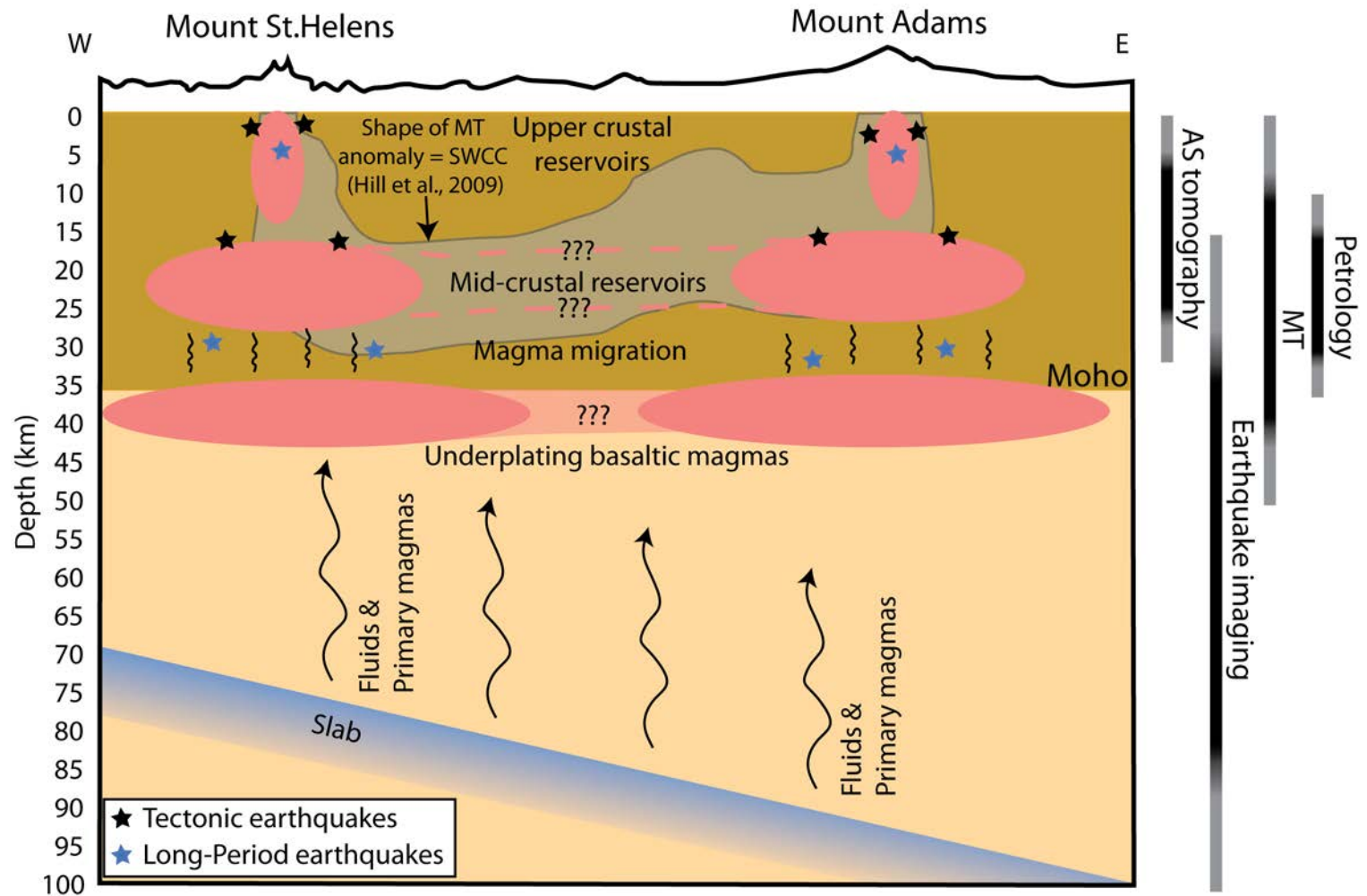
- Geophysics
 - Active-source seismology
 - Levander (Rice) and collaborators
 - Earthquake seismology
 - Abers (LDEO), Creager, Vidale, Houston (UW), Moran, Denlinger (USGS), Levander (Rice)
 - Electromagnetic Imaging (MT)
 - Schultz (OSU), Bedrosian (USGS)
- Petrology – magma chamber dynamics
 - Sisson, Clynne, Pallister (USGS), Bachmann (UW),

Why Mt St. Helens?

- It's active!
- Reasonable size
- Well-characterized
- Access for instrumentation



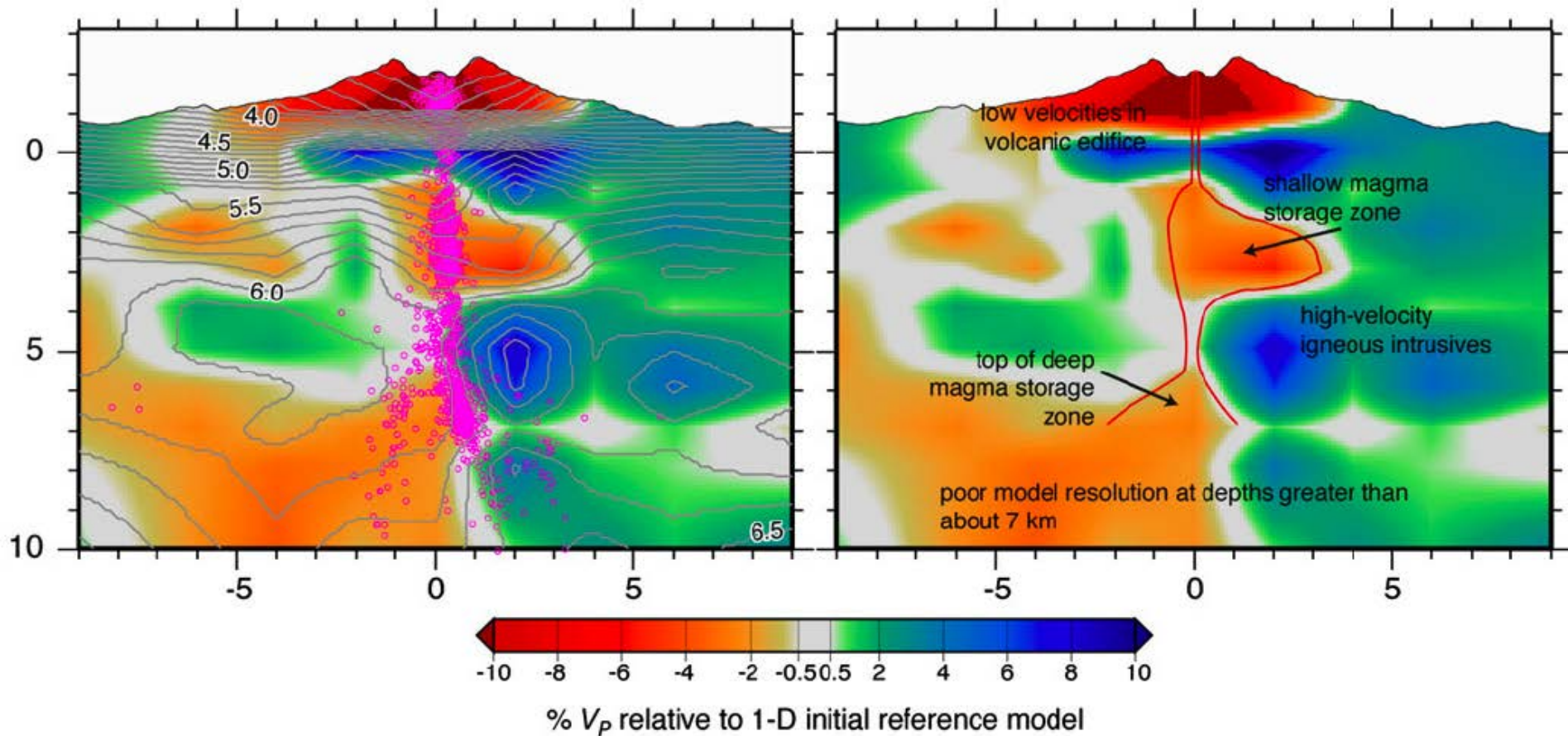
Imaging Targets



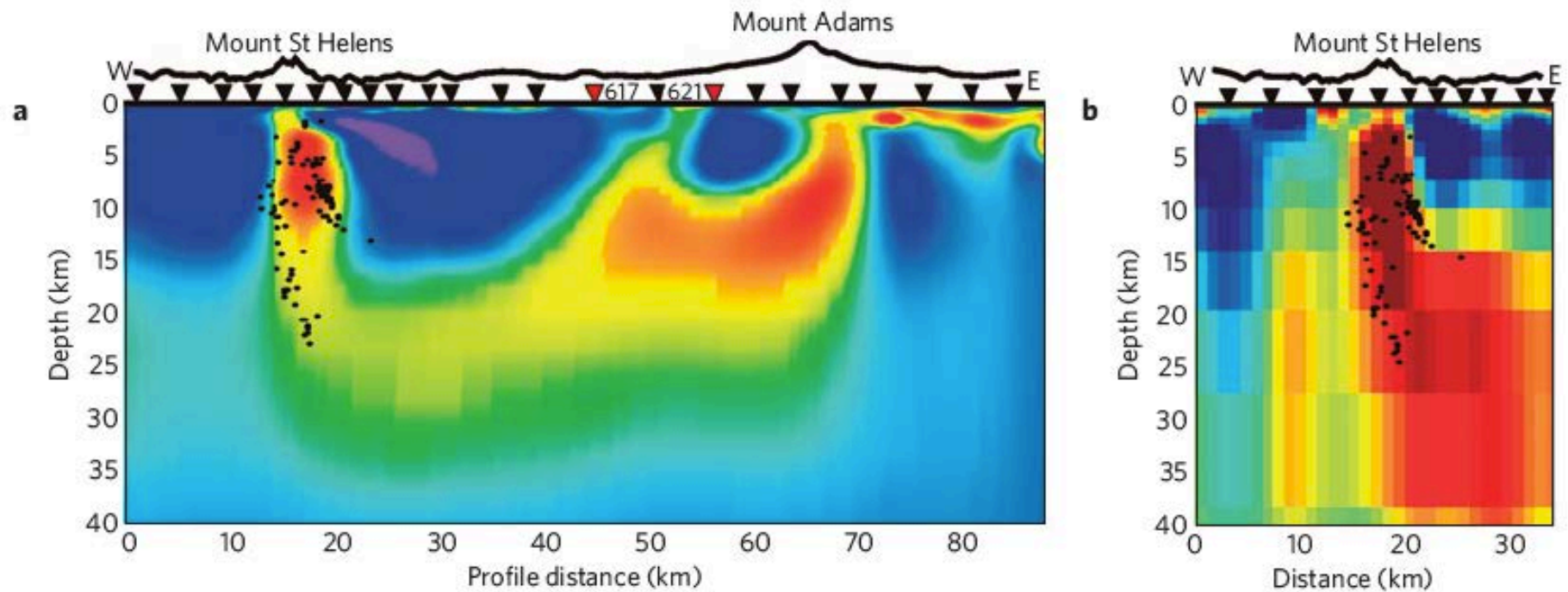
Some existing imaging

Upper crustal magma reservoir

E-W cross-section

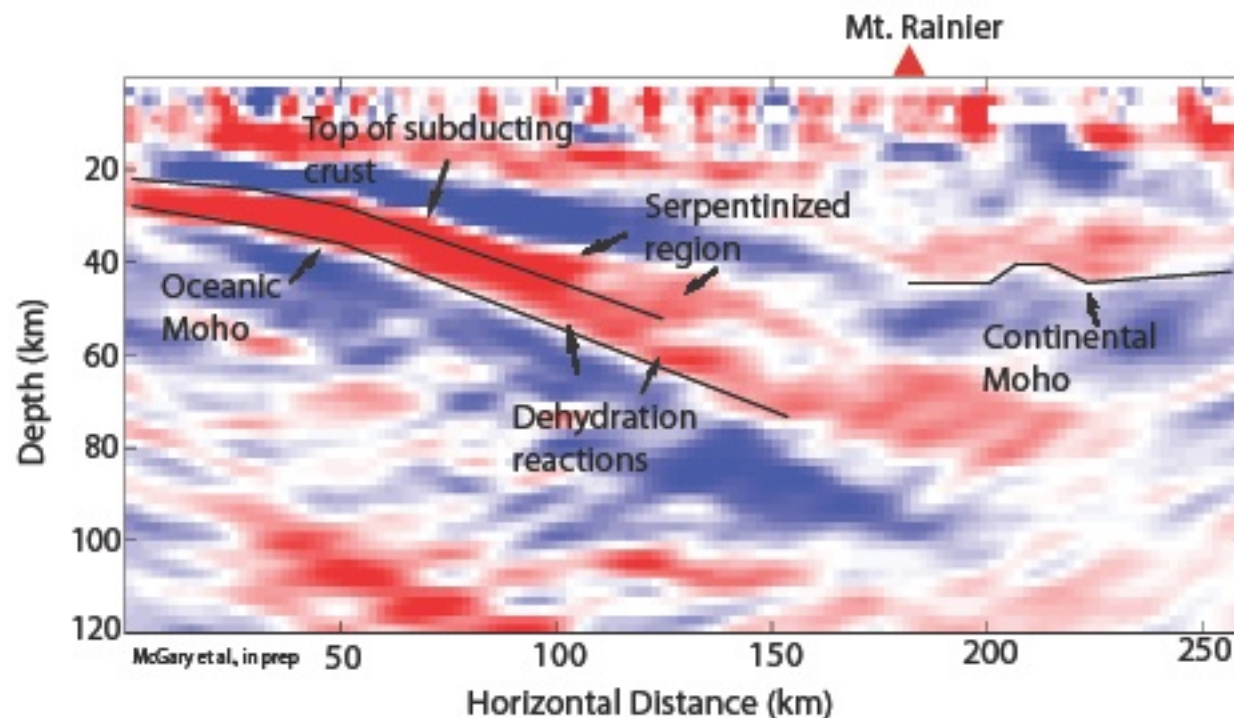
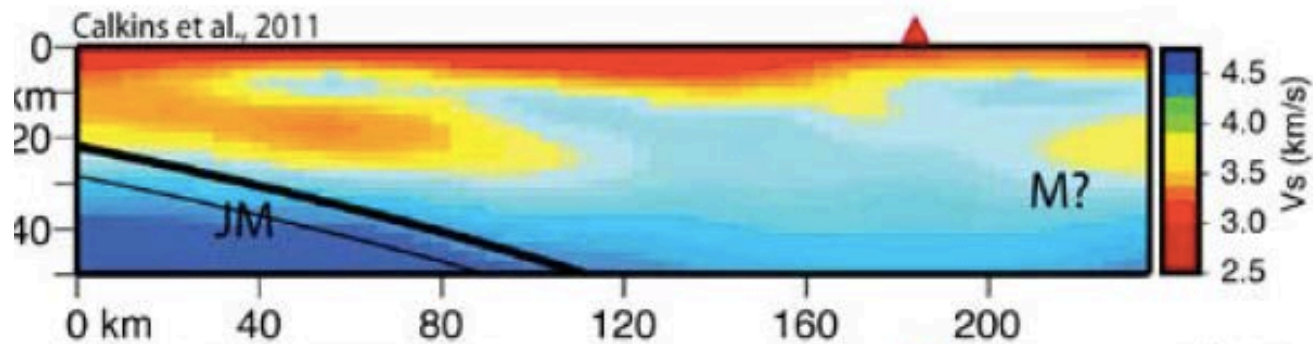


Magma plumbing from existing MT

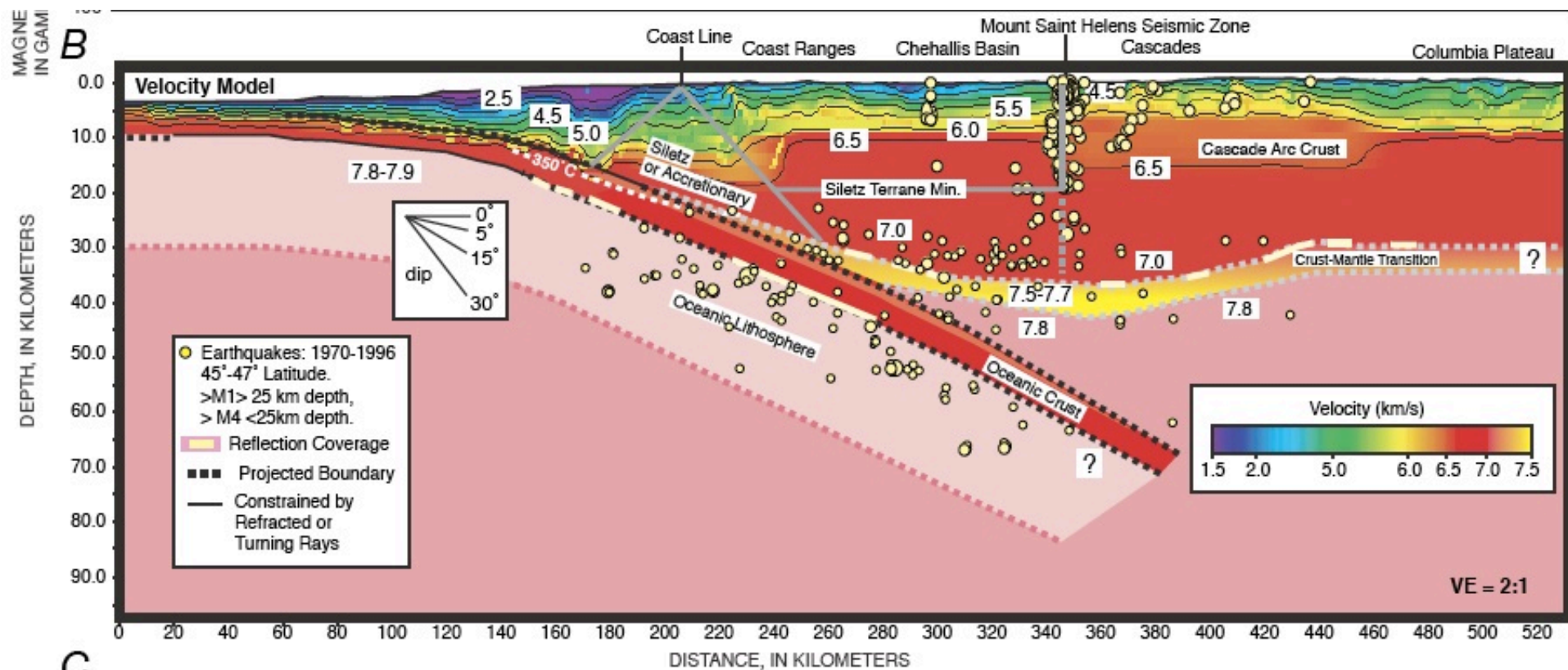


Hill et al., 2009

Existing seismic imaging

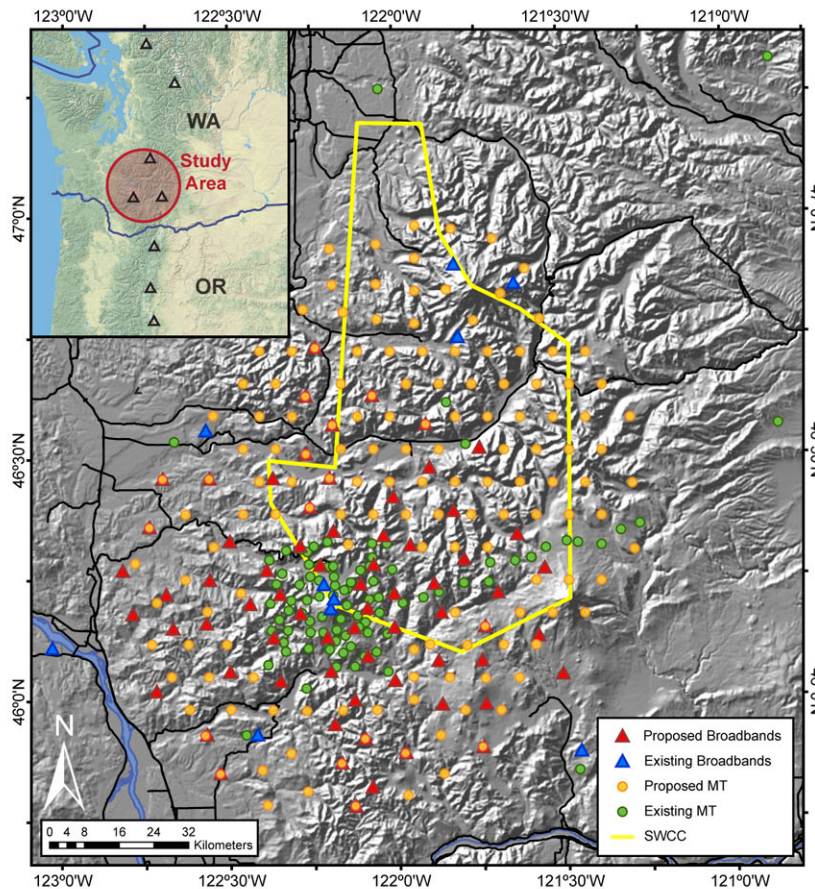


Seismic velocity model

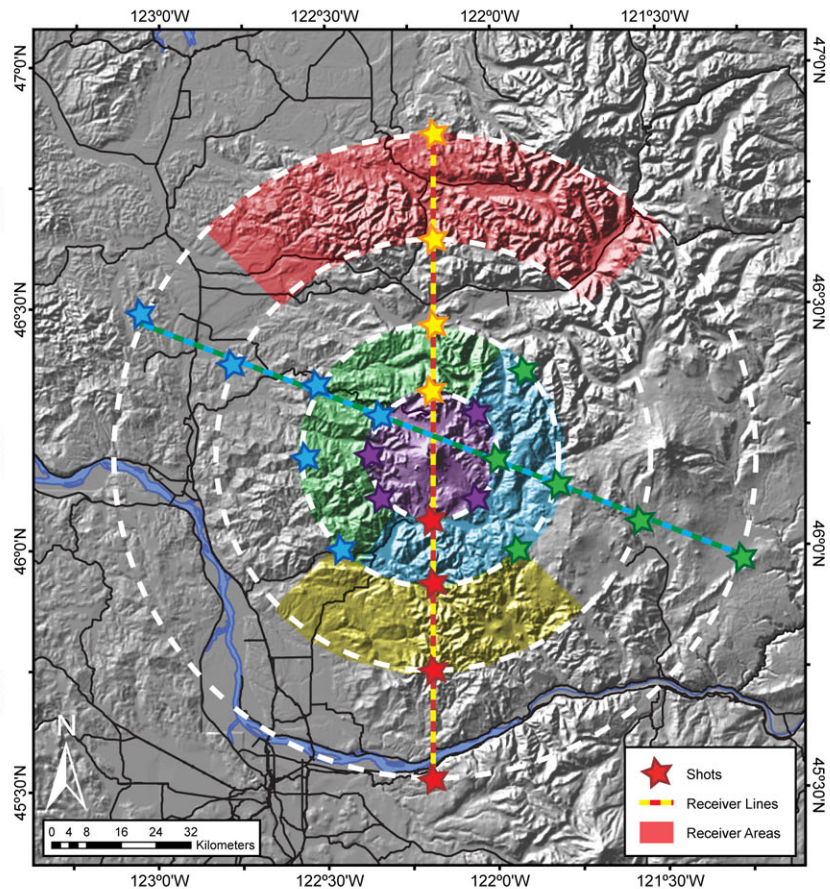


Parsons et al., 1998

Planned arrays



Passive Seis: 70 broadband + existing
MT: 150 wideband sites



Active 3D Seis: 2600 Texans/multiple deploy



Some Impacts

- Long term goal:
 - Try to see magmas from their mantle source to the near-surface
- Near term goal:
 - When complete, we will produce the largest, most comprehensive dataset for plumbing of any volcanic system
- Vehicle for publicizing GeoPRISMS science broadly
- Strong collaborations with other groups